

A Rare Opportunity, the Mu2e Experiment

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Fermilab

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Topics I'll Cover Today

- What is High Energy Physics?
- What is a muon?
- What is “Mu2e”?
... and how does it work?
... and why is it important?

Preamble

What is Science?

What is Science?

- Evidence based

For first few millenia

- There were no scientists... only “Natural Philosophers”



Birth of Modern Science

- Galileo Galilei insists that scientific conclusions must be based on experimental evidence



Galileo Galilei, 1636

What is Science?

- Evidence based

What is Science?

- Evidence based
- Curiosity driven



What is High Energy Physics?

What is High Energy Physics?



It is a branch of science

What is the goal of High Energy Physics?

Identify the fundamental particles
...and the forces that govern their interactions

What is the goal of High Energy Physics?

What are things made of?

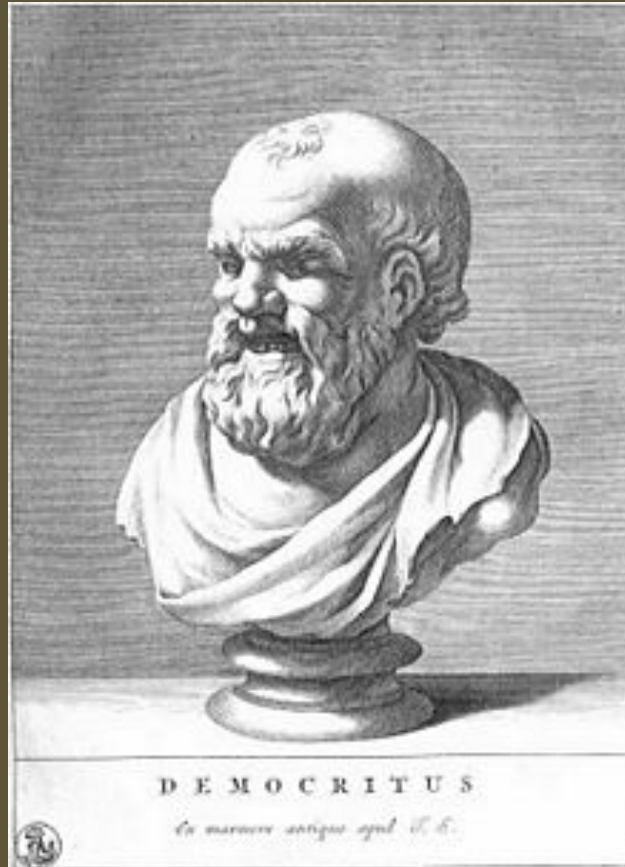
Why do they behave the way they do?

Our first guess...



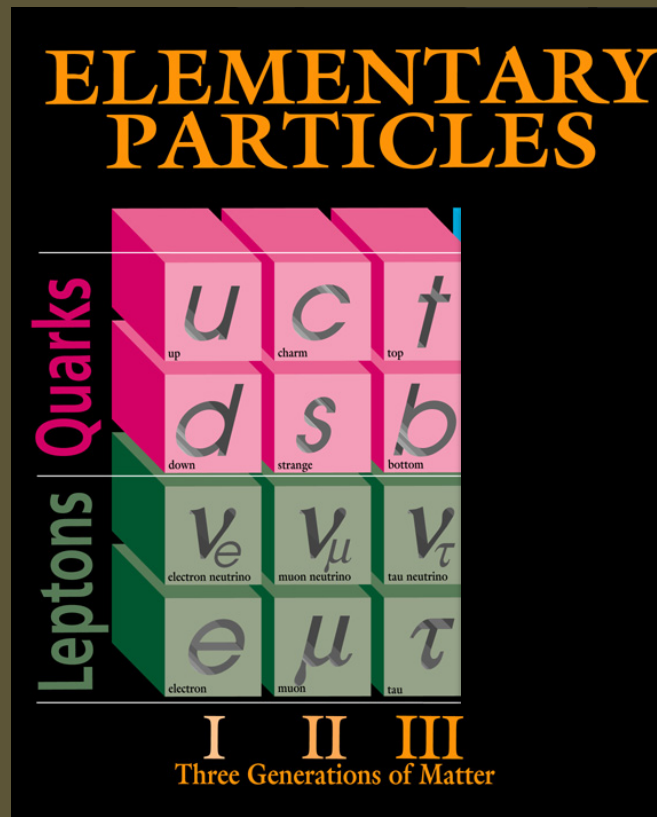
earth, wind, fire, water

A couple thousand years later...



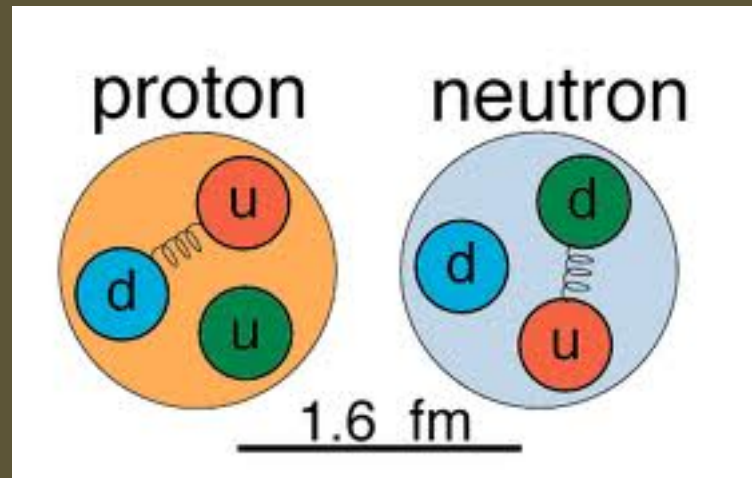
Postulated that all things are made of “atoms”

Today



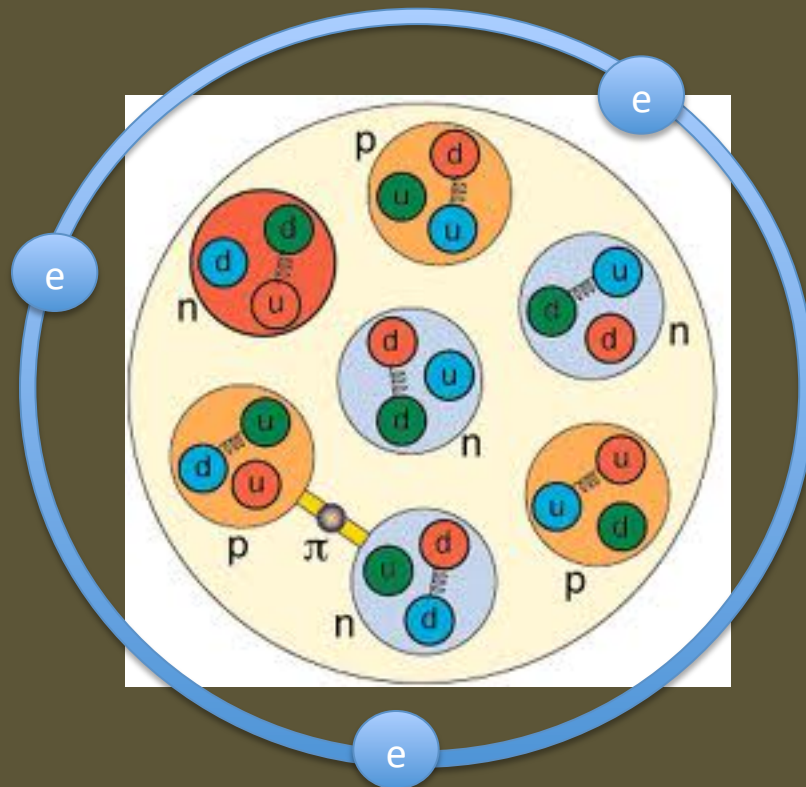
The fundamental particles

From Quarks to Cosmos



At its most basic level, everything we see is
made of quarks and electrons

From Quarks to Cosmos



At its most basic level, everything we see is
made of quarks and electrons

From Quarks to Cosmos

Periodic Table of the Elements

1A	2A	3A	4A	5A	6A	7A	8A																								
1 H							2 He																								
3 Li	4 Be						10 Ne																								
11 Na	12 Mg	13 Al	14 Si	15 P	16 S	17 Cl	18 Ar																								
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr														
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe														
55 Cs	56 Ba	57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
87 Fr	88 Ra	89 Ac	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og

• Lanthanide Series
+ Actinide Series

At its most basic level, everything we see is made of quarks and electrons

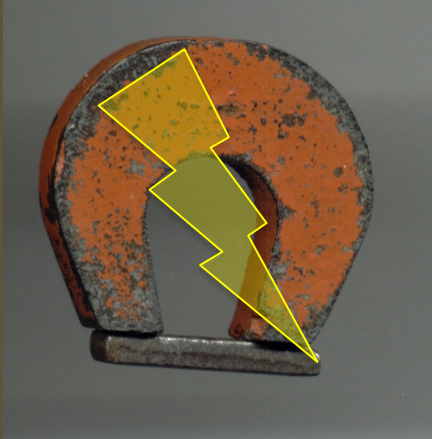
From Quarks to Cosmos



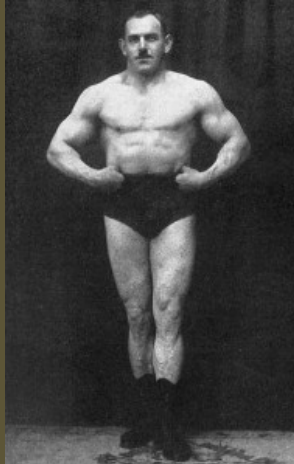
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Today

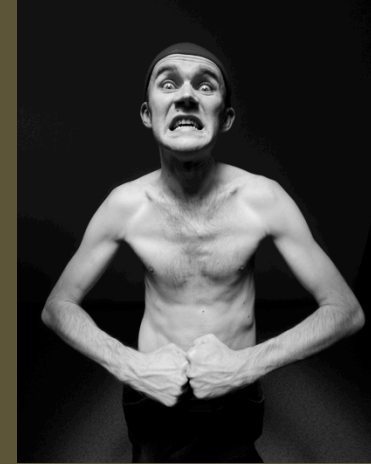
electromagnetic



strong



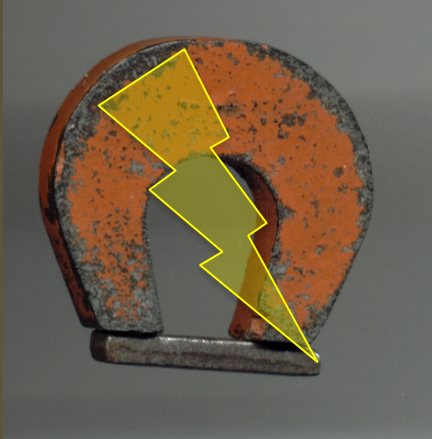
weak



The forces that govern their interactions

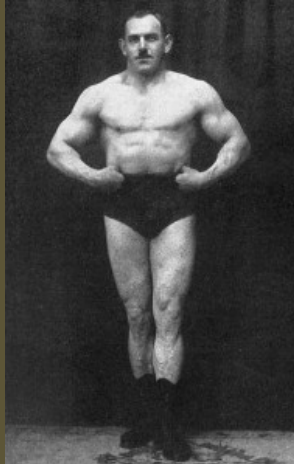
Today

electromagnetic



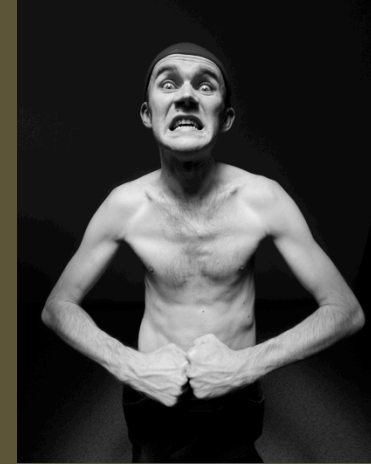
γ

strong



g

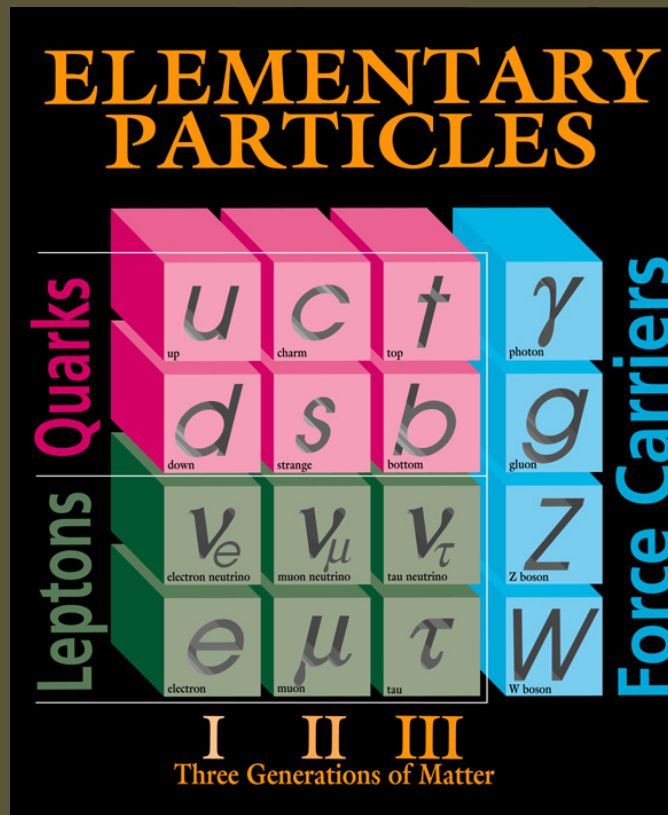
weak



W, Z

The forces that govern their interactions

Today



The Standard Model

Standard Model Evidence

- Discovery of the γ



Standard Model Evidence

- Discovery of the γ
- Discovery of the nucleus



Standard Model Evidence

- Discovery of the γ
- Discovery of the nucleus
- Discovery of the μ



Standard Model Evidence

- Discovery of the γ
- Discovery of the nucleus
- Discovery of the μ
- Discovery of the ν



Standard Model Evidence

- Discovery of the γ
- Discovery of the nucleus
- Discovery of the μ
- Discovery of the ν
- Discovery of c quark



Standard Model Evidence

- Discovery of the γ
- Discovery of the nucleus
- Discovery of the μ
- Discovery of the ν
- Discovery of c quark
- Discovery of anti-matter

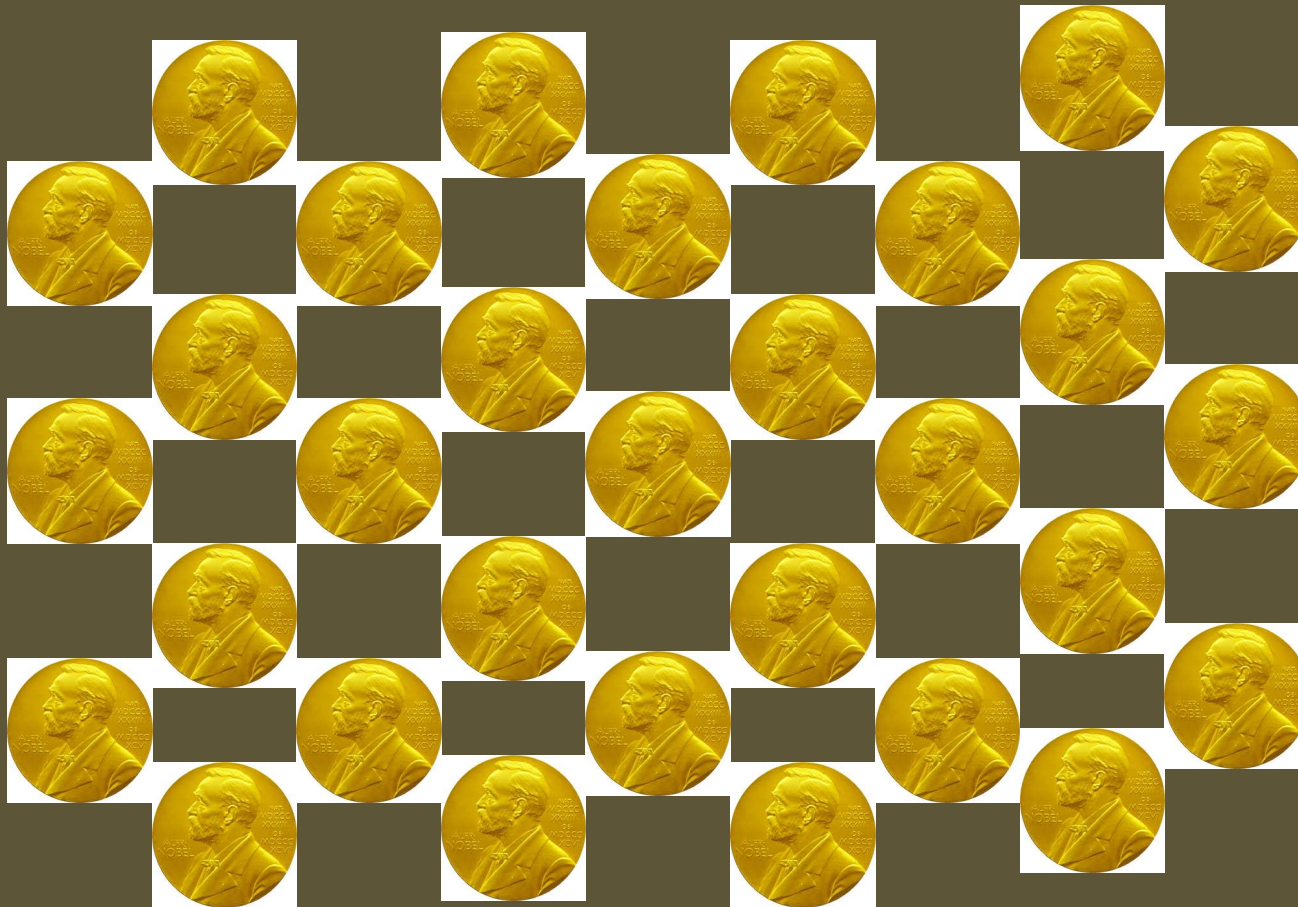


Standard Model Evidence

- Discovery of the γ
- Discovery of the nucleus
- Discovery of the μ
- Discovery of the ν
- Discovery of c quark
- Discovery of anti-matter
- Discovery of W & Z



Standard Model Evidence



... and 1000s of additional experimental results

But there is a catch



We know the Standard Model is incomplete



An (obvious) missing piece...

gravity

21-Mar-2012

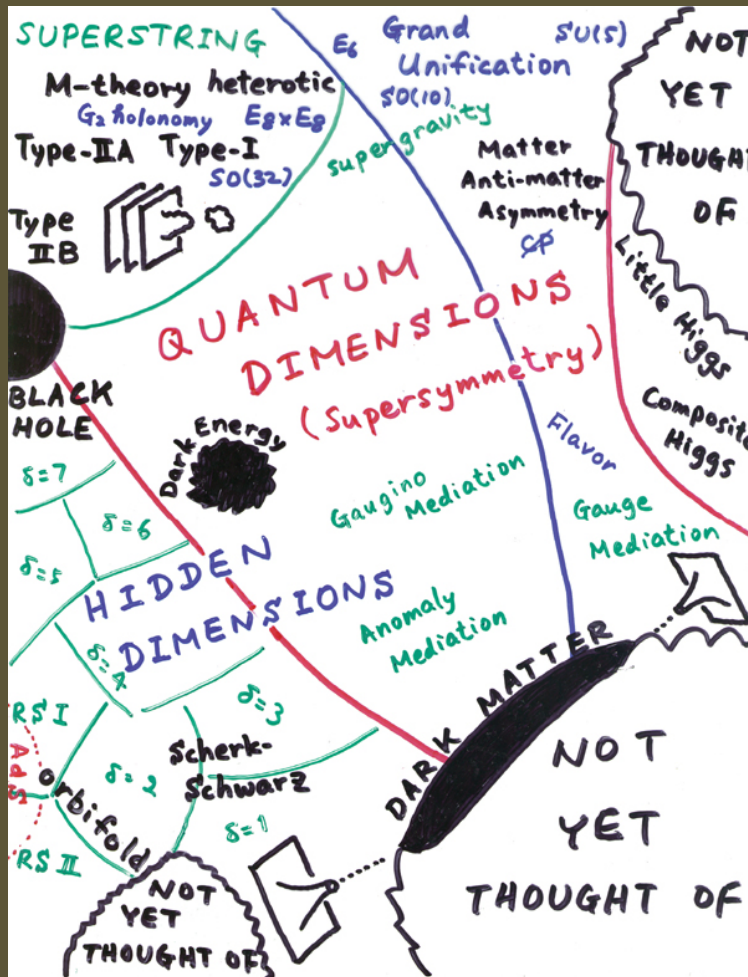
D.Glenzinski, Fermilab

There is a catch



Standard Model only describes 5% of Universe

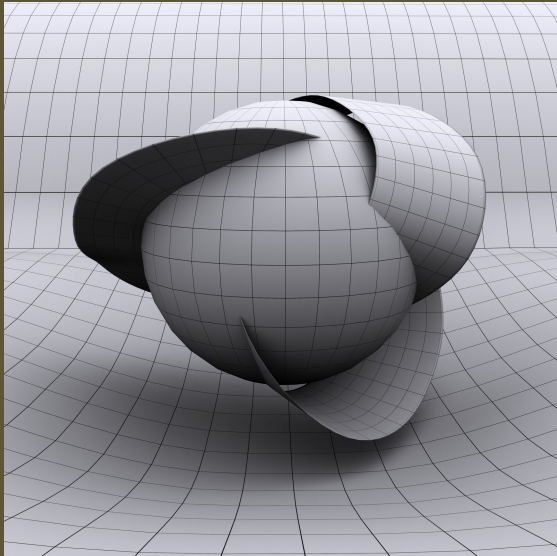
Possibilities



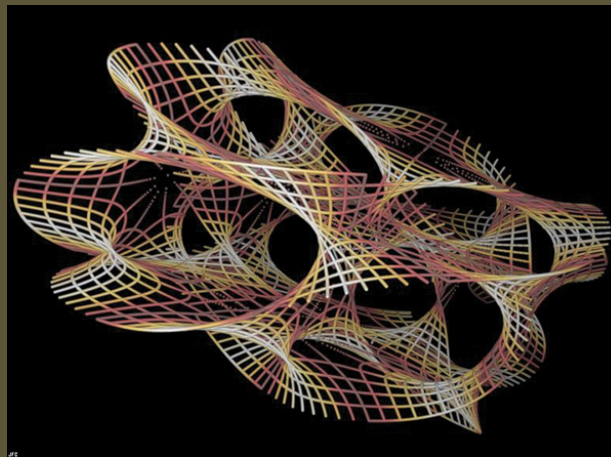
We have lots of ideas of what a more complete theory might look like...

but we don't know which one (if any) is correct.

In the future



Extra Dimensions



Super Strings

ELEMENTARY PARTICLES

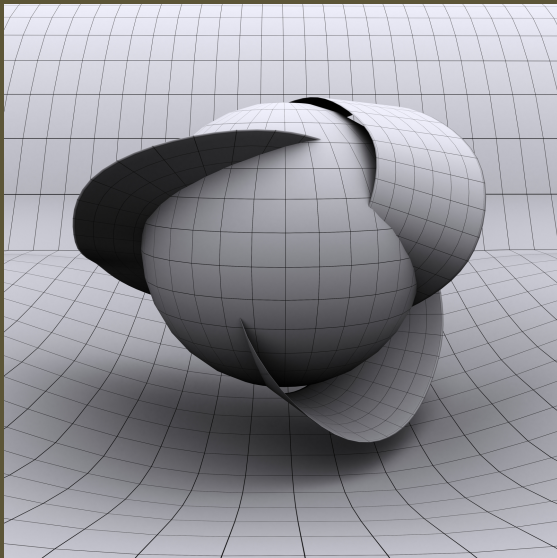
Quarks	u up	c charm	t top	γ photon
	d down	s strange	b bottom	g gluon
Leptons	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z Z boson
	e electron	μ muon	τ tau	W W boson
				Force Carriers

Sleptons	\tilde{u} up	\tilde{c} charm	\tilde{t} top	$\tilde{\gamma}$ photon
	\tilde{d} down	\tilde{s} strange	\tilde{b} bottom	\tilde{g} gluon
Squarks	$\tilde{\nu}_e$ electron neutrino	$\tilde{\nu}_\mu$ muon neutrino	$\tilde{\nu}_\tau$ tau neutrino	χ^0
	\tilde{e} electron	$\tilde{\mu}$ muon	$\tilde{\tau}$ tau	χ^-
				Gauginos

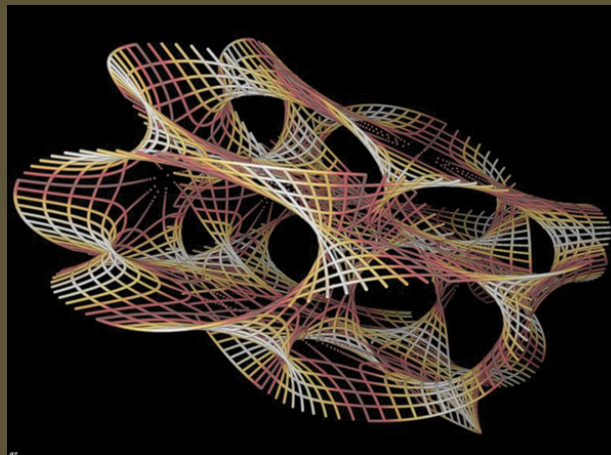
Super Symmetry

Our picture of HEP may well be very different.

In the future



Extra Dimensions



Super Strings

ELEMENTARY PARTICLES

Quarks	u up	c charm	t top	γ photon
	d down	s strange	b bottom	g gluon
Leptons	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z Z boson
	e electron	μ muon	τ tau	W W boson
				Force Carriers

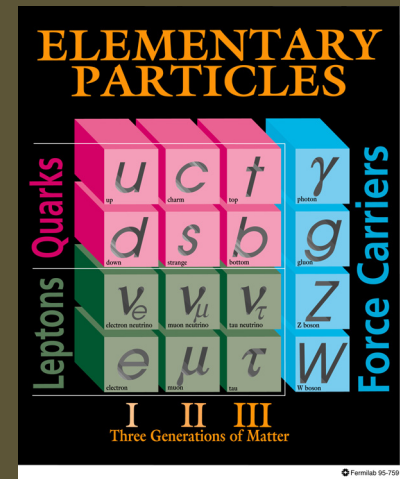
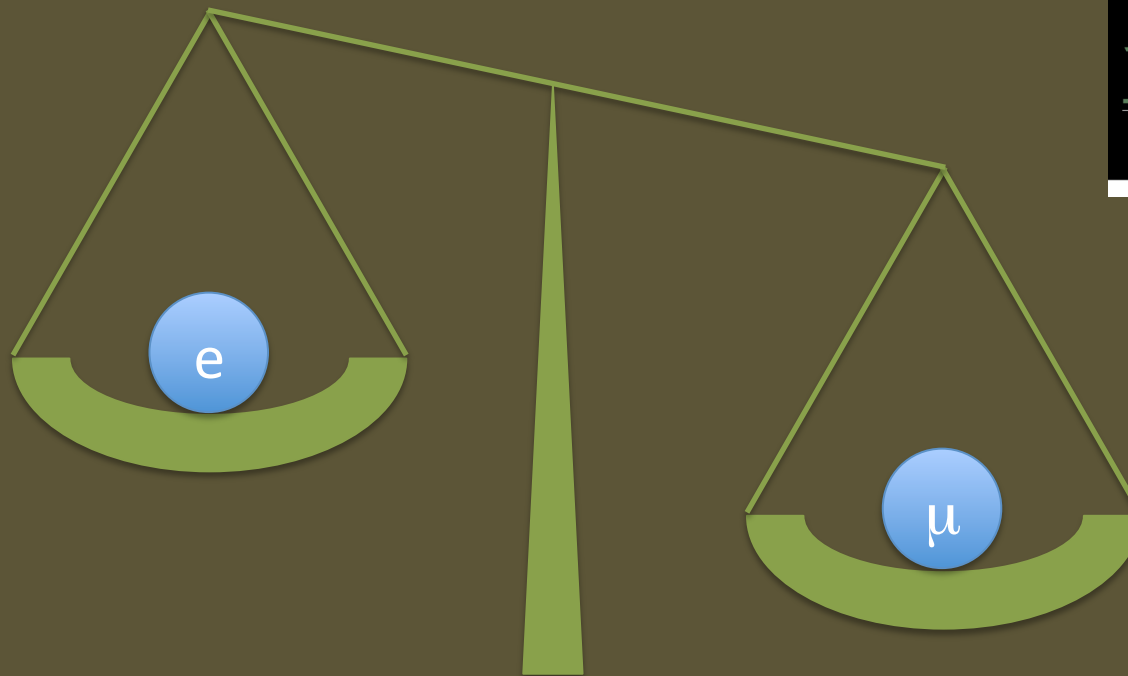
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Squarks	$\tilde{\nu}_e$ electron neutrino	$\tilde{\nu}_\mu$ muon neutrino	$\tilde{\nu}_\tau$ tau neutrino	χ^0
	\tilde{e} electron	$\tilde{\mu}$ muon	$\tilde{\tau}$ tau	χ^-
				Gauginos

Super Symmetry

Our picture of HEP may well be very different.
We need experiments (like Mu2e) to point the way.

What is a muon?

What is a muon?



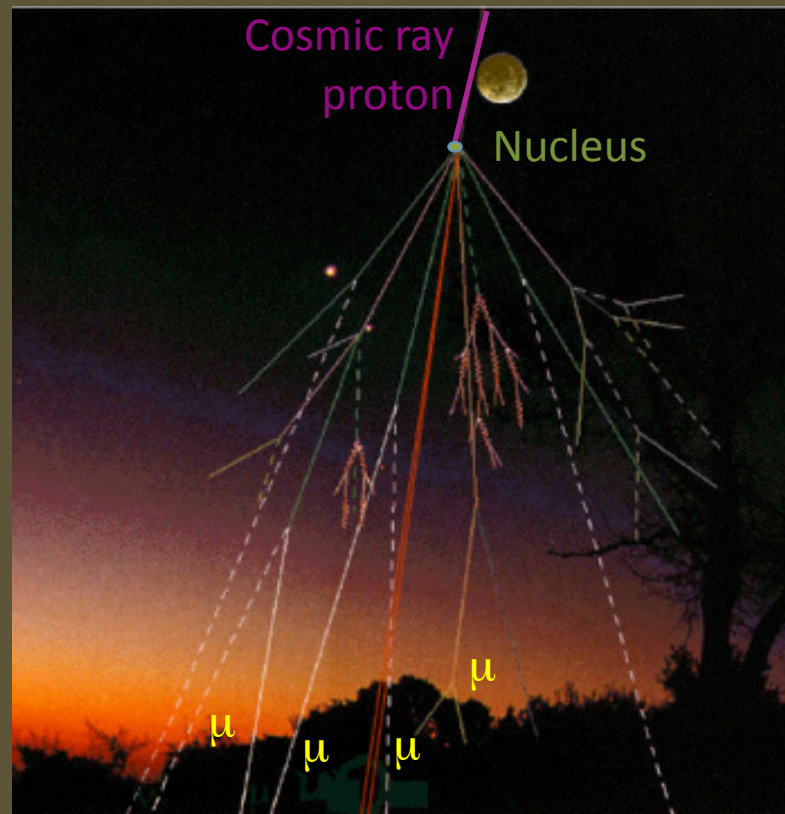
Very much like an electron... but heavier

What is a muon?



discovered in 1936 using cosmic rays
(they are streaming through us now)

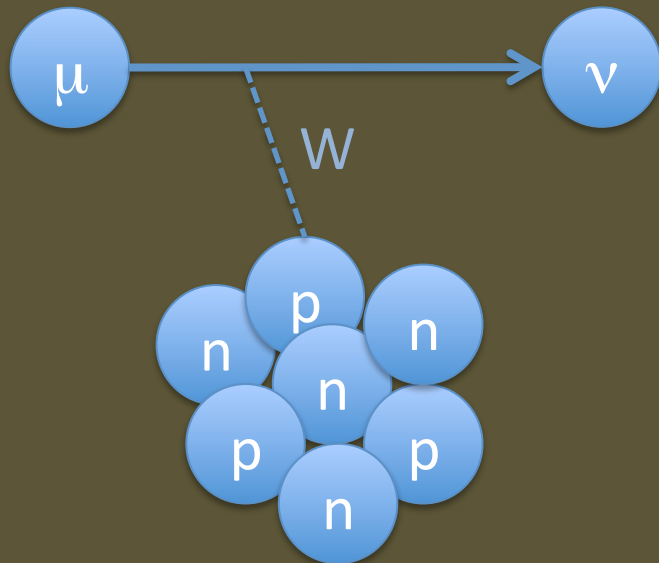
What is a muon?



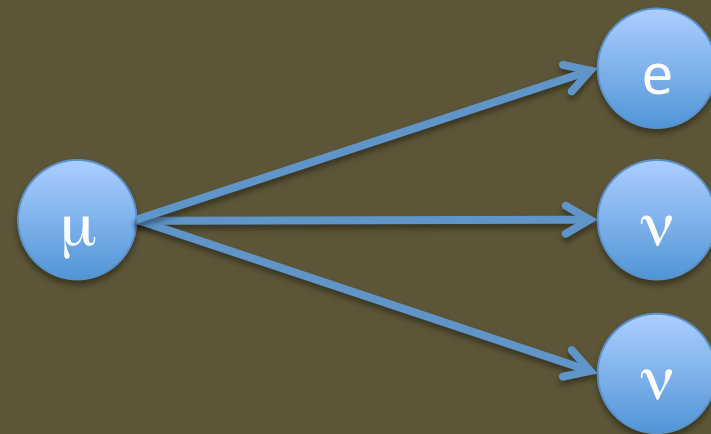
discovered in 1936 using cosmic rays
(they are streaming through us now)

What is a muon?

- μ s have only ever been observed to do one of two things...



1) Interact with a nucleus
to produce a ν



2) Decay into an electron
and two ν

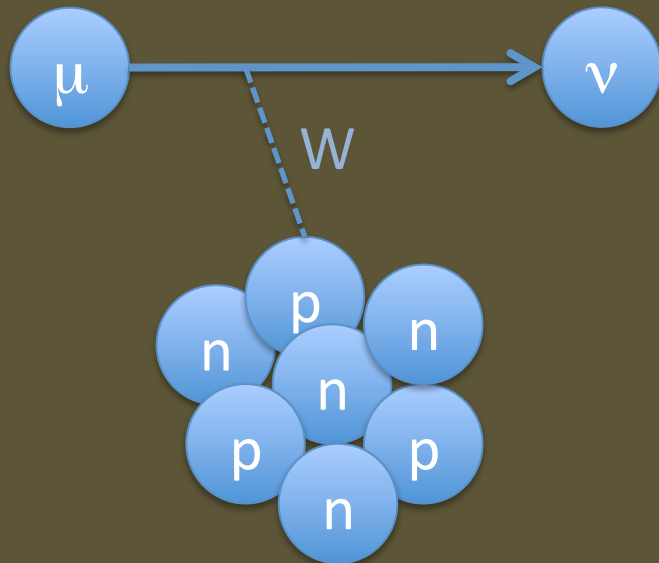
What is “Mu2e”?

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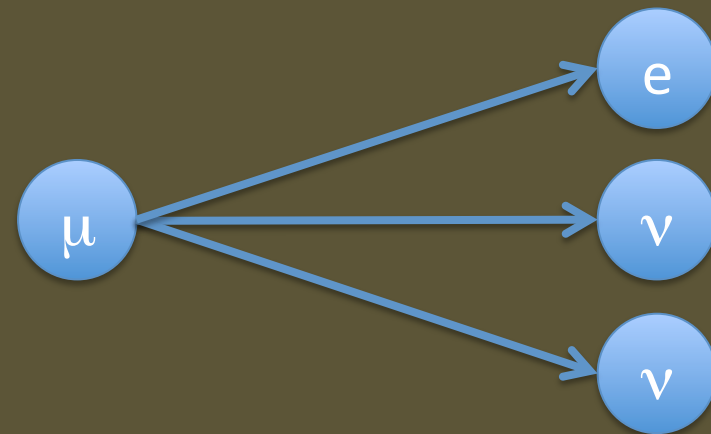
- A high energy physics experiment
- Uses μ_s to look for a very rare process

What is Mu2e?

- μ s have only ever been observed to do one of two things...



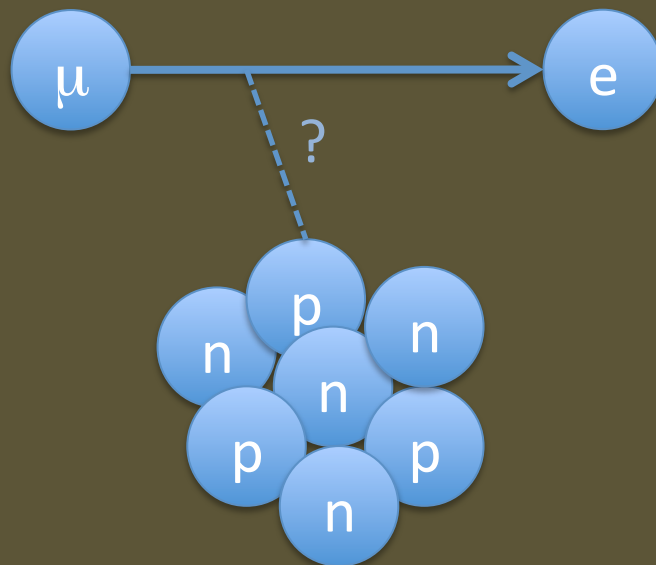
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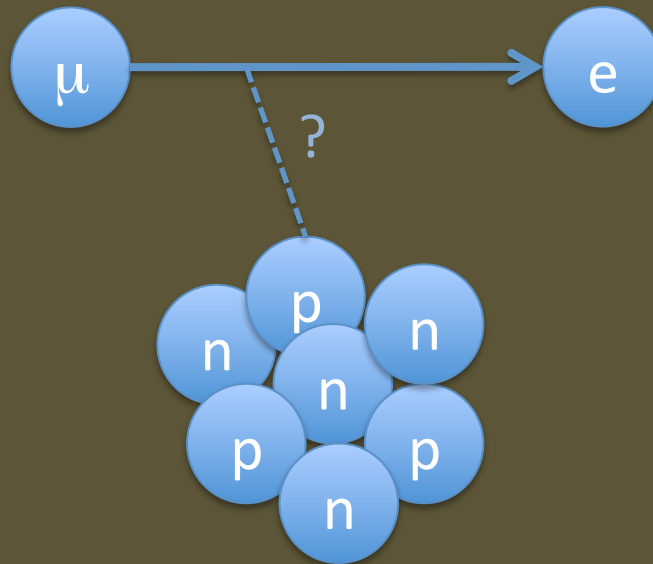
What is Mu2e?

- Mu2e is looking for evidence of a 3rd thing...



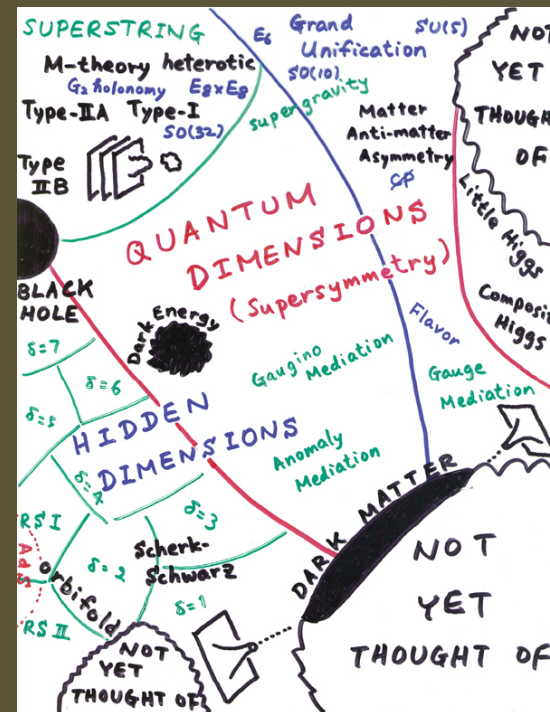
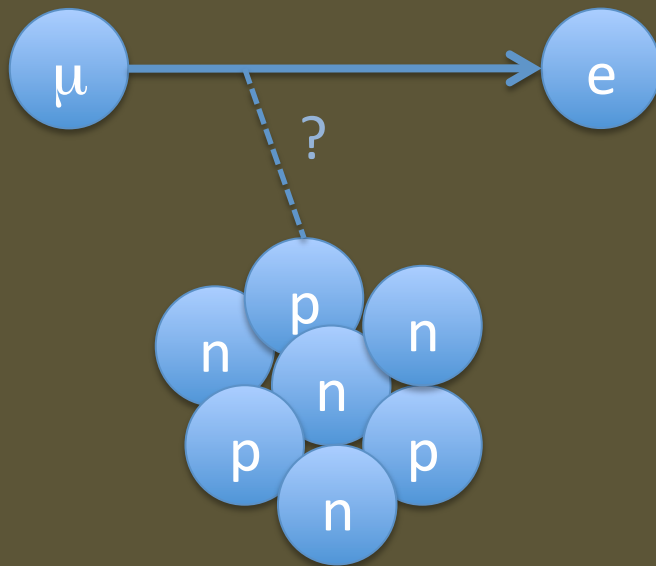
3) Interact with a nucleus to produce an electron

Why is Mu2e important?



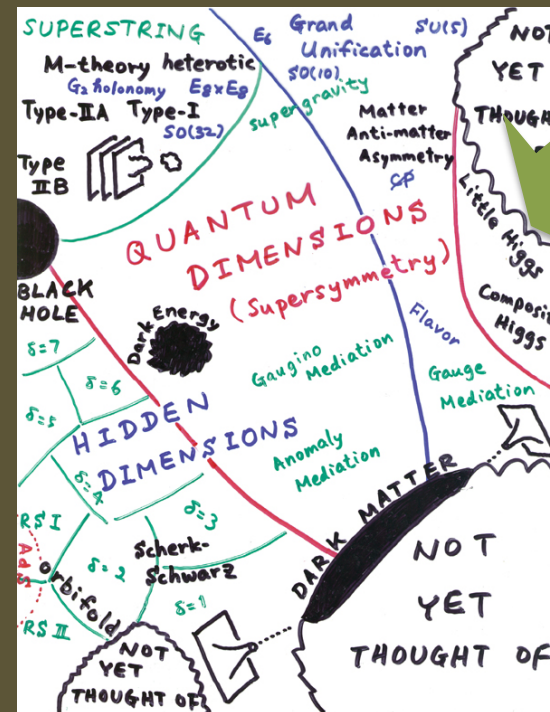
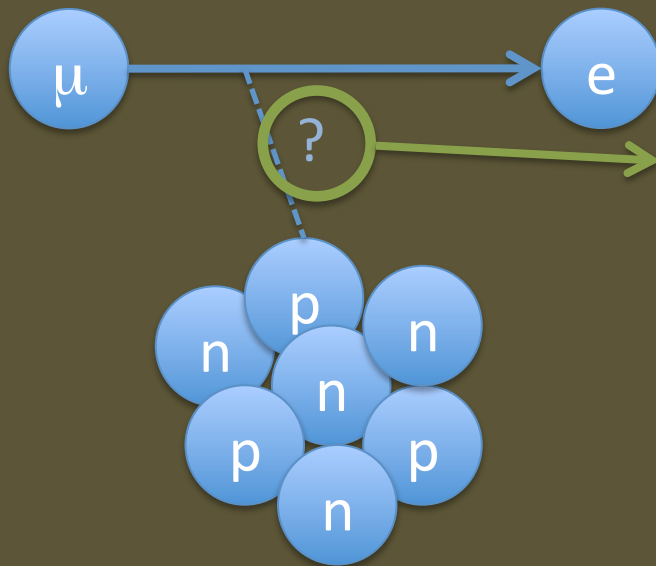
There are many ways this process can occur...
but *none* of them are in the Standard Model!

Why is $\mu \rightarrow e$ important?



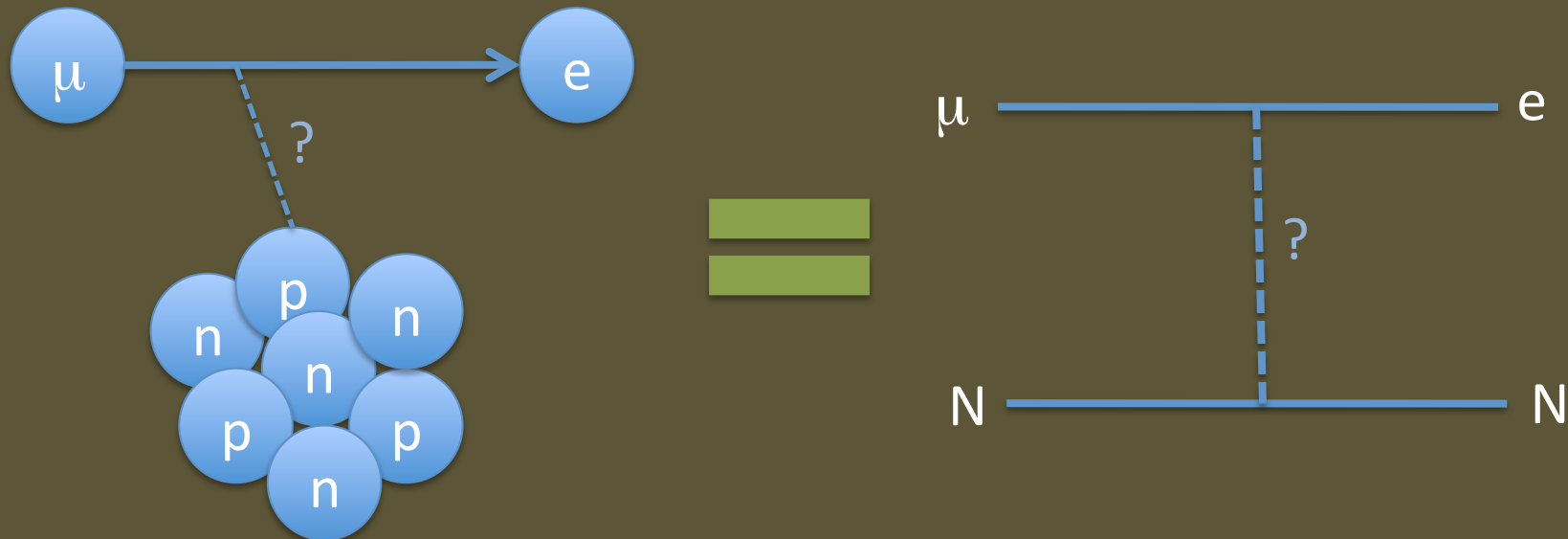
So, by measuring the rate of the $\mu N \rightarrow e N$ process we can test these new theories

Why is Mu2e important?



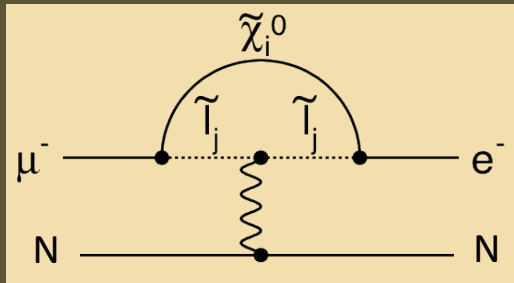
The rate tells you something about the question mark in the figure, which in turn depends on the details of the new theory

Why is Mu2e important?

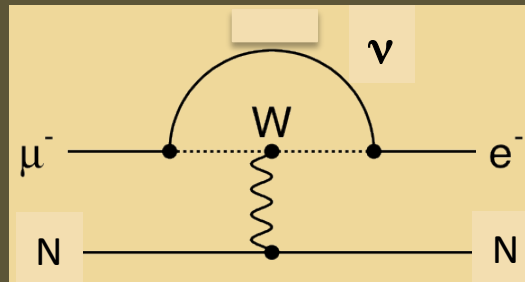


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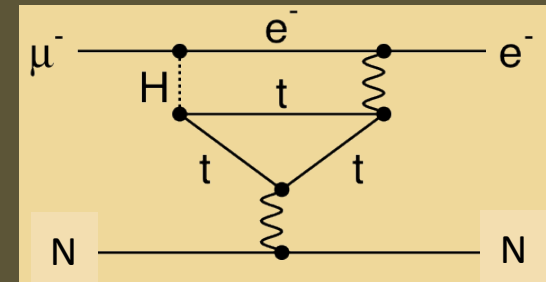
Why is Mu2e important?



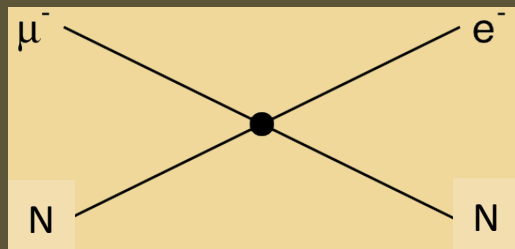
Supersymmetry



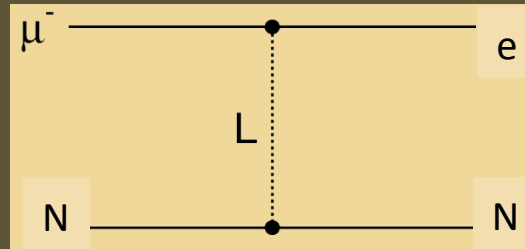
Heavy Neutrinos



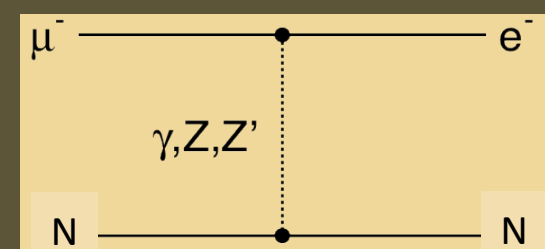
Two Higgs Doublets



Preons



Leptoquarks



New Heavy Bosons
New Forces

because Mu2e will be sensitive to many theories
of “physics beyond the Standard Model”

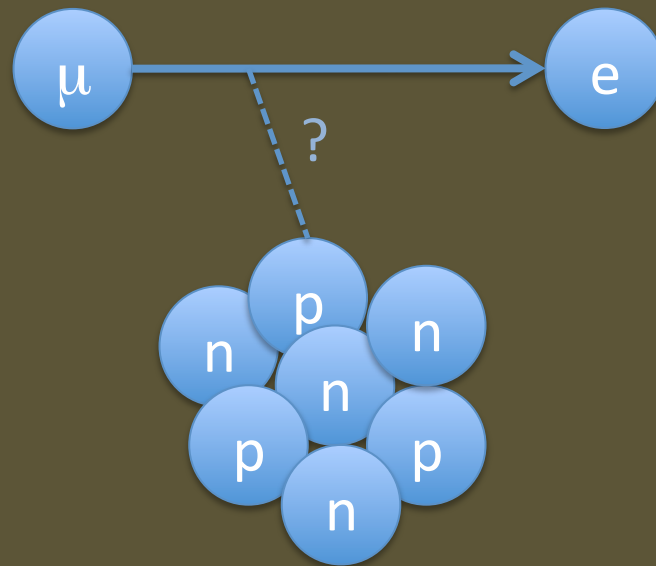
Physicist Speak Decoded

“sensitive to
physics
beyond the
Standard
Model”



“can make
Major
Discovery”

Why is Mu2e important?



because Mu2e is poised to make a major discovery

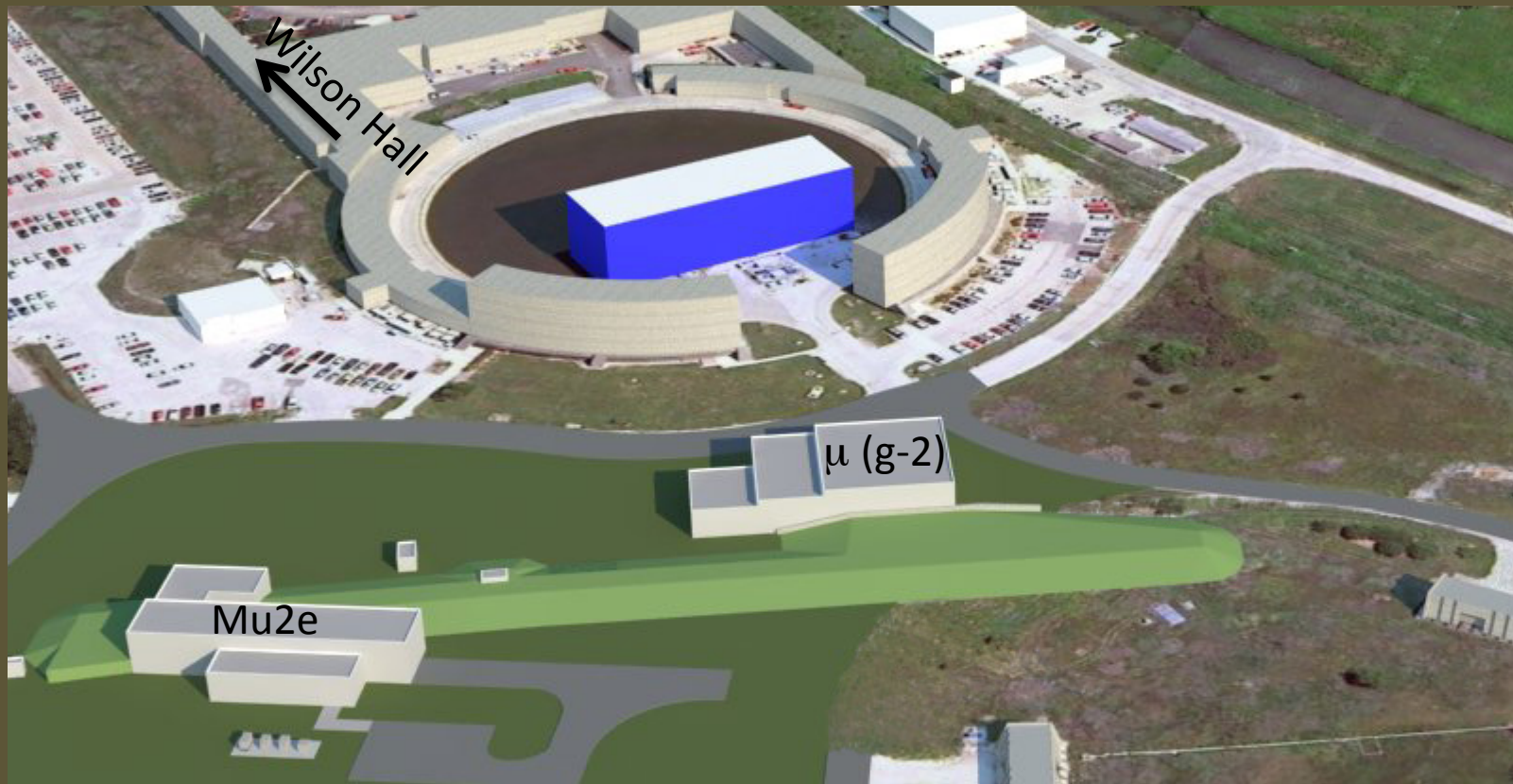
How does Mu2e work?

How does Mu2e work?



Mu2e will recycle some of the accelerator components used by the Tevatron to make anti-protons and instead make μ_s

How does Mu2e work?



Mu2e will be located at a new “Muon Campus”

Mu2e Building

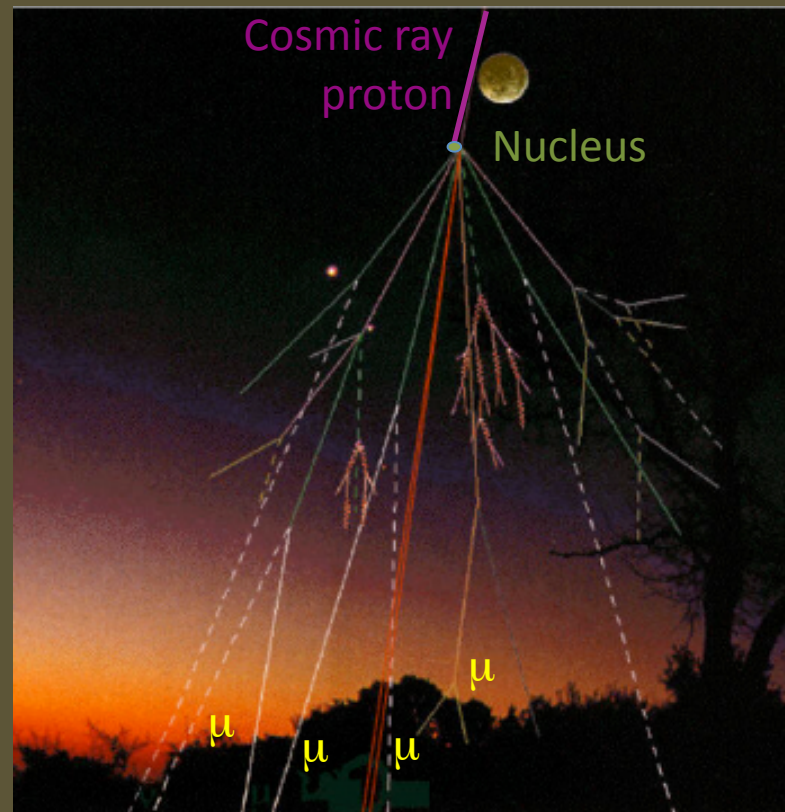


Graphic of proposed Mu2e detector hall

How does Mu2e work?

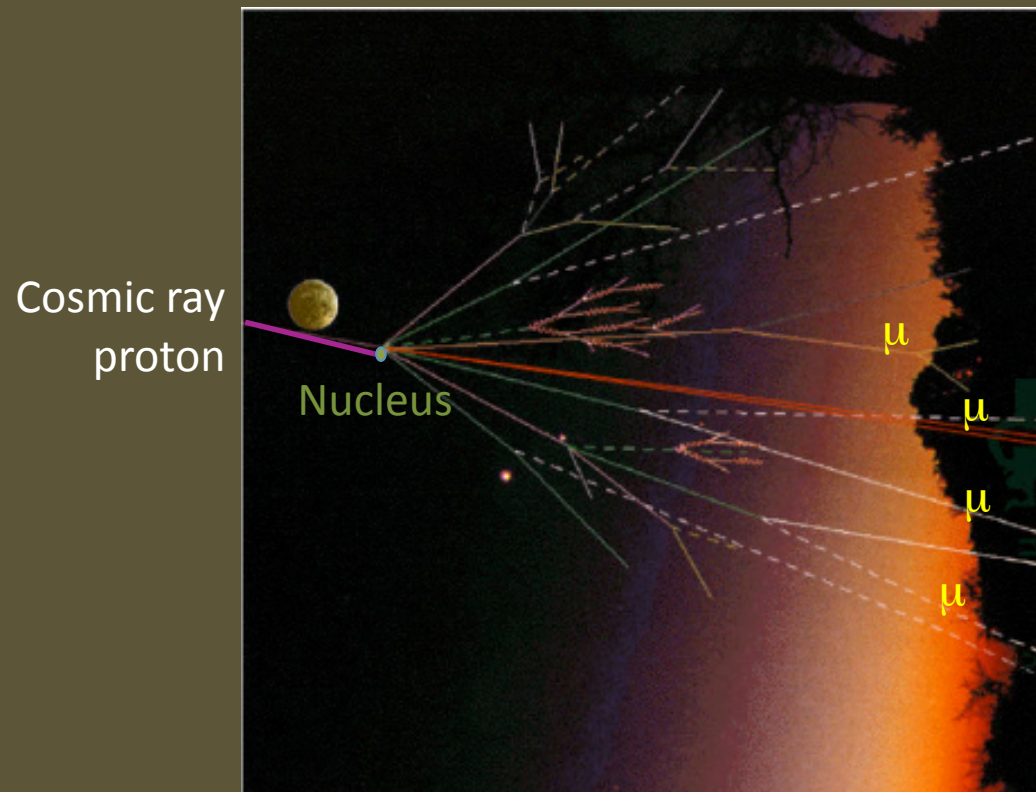
- 1) Make lots of μ s
- 2) Stop them near nuclei
- 3) Look for electrons

Making μ s



smash protons on nuclei, collect the debris
(just like cosmic rays)

Making μ_s

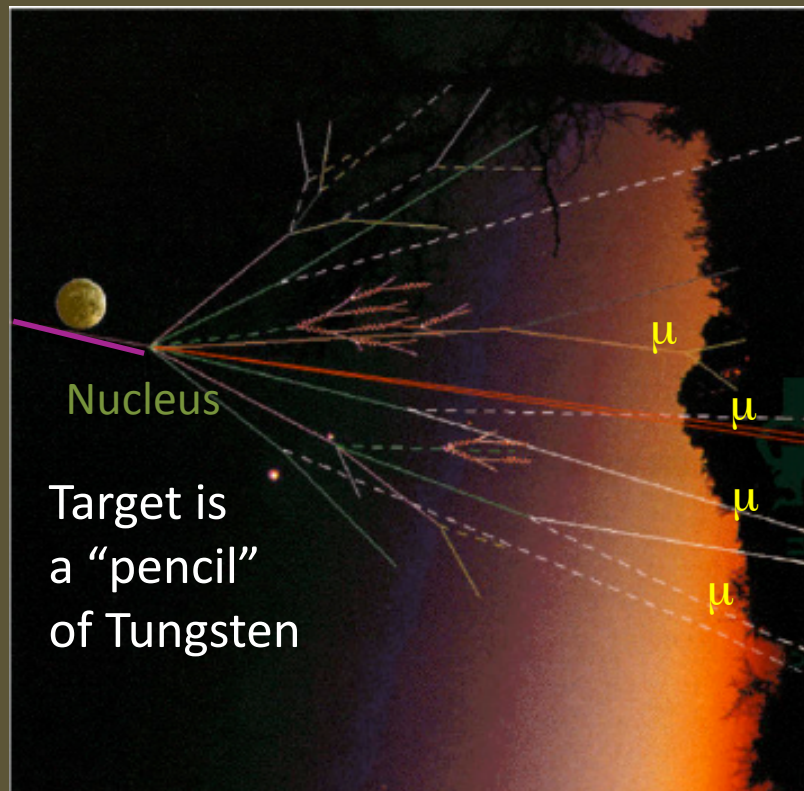


smash protons on nuclei, collect the debris
(just like cosmic rays)

Making μ s

Protons from Booster

- 8 GeV kinetic energy
- 99.4% speed of light
- 667,000,000 mph

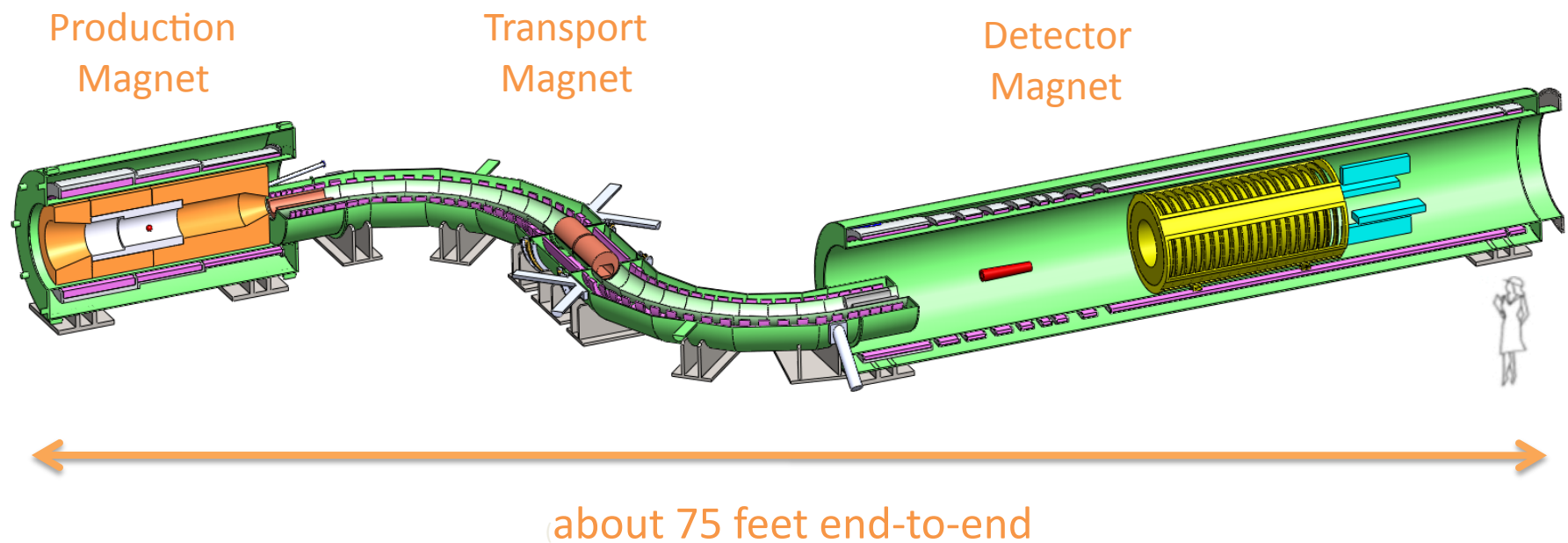


The debris eventually decays, often to μ s

Use magnets to collect the μ s

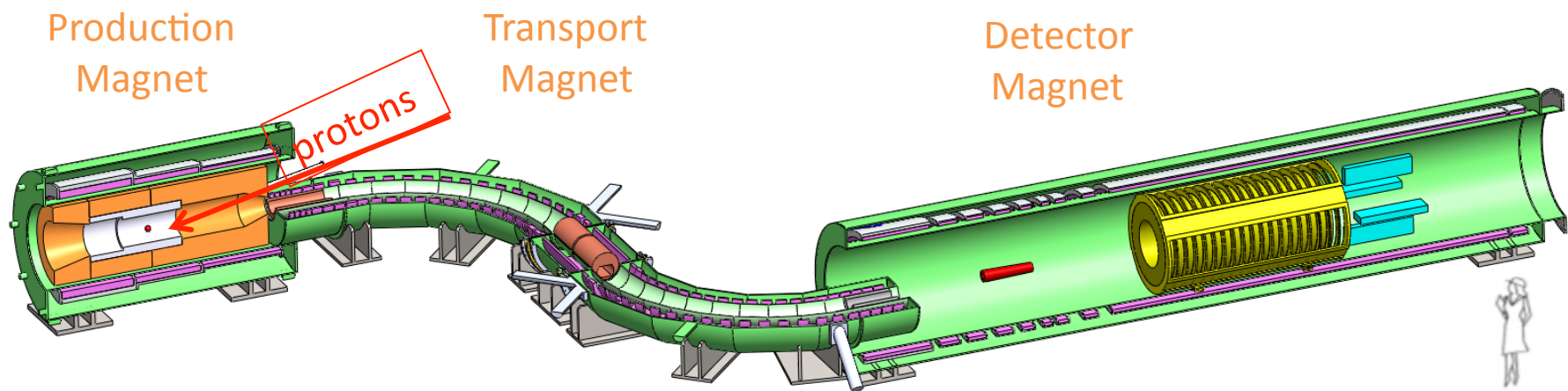
smash protons on nuclei, collect the debris
(just like cosmic rays)

Mu2e Apparatus



- Mu2e will consist of 3 magnet systems

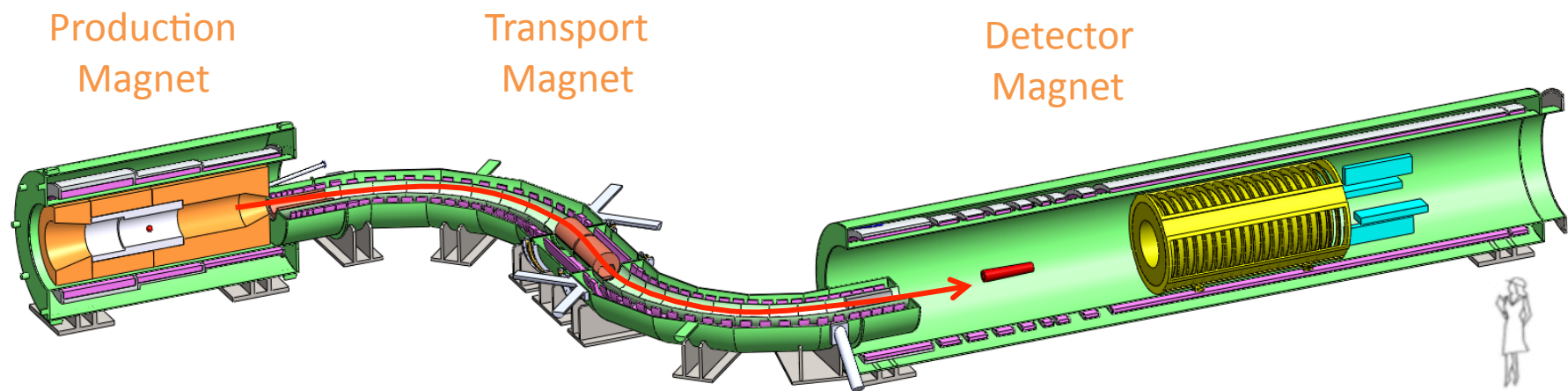
Mu2e Apparatus



Production Magnet:
this is where the μ s get made

- Mu2e will consist of 3 magnet systems

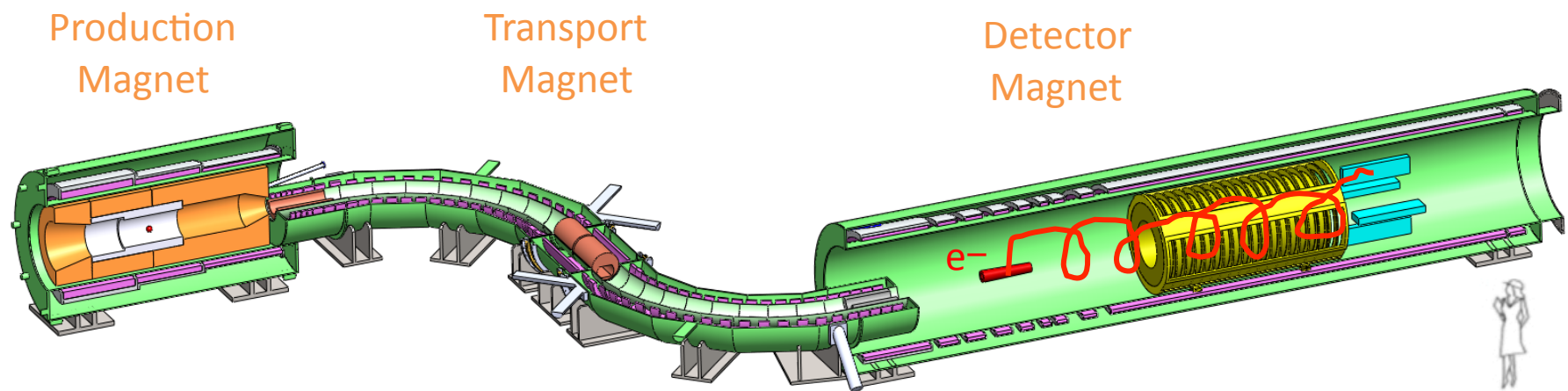
Mu2e Apparatus



Transport Magnet:
“carries” the μ s to the Detector Magnet

- Mu2e will consist of 3 magnet systems

Mu2e Apparatus



Detector Magnet:
this is where we look for evidence $\mu N \rightarrow e N$

- Mu2e will consist of 3 magnet systems

Some Mu2e numbers

- Every 1 second Mu2e will
 - Send 8,000,000,000,000 protons to the Production Magnet
 - Send 26,000,000,000 μ_s through the Transport Magnet
 - Stop 13,000,000,000, μ_s in the Detector Magnet
- By the time Mu2e is done, about 1,000,000,000,000,000,000 stopped μ_s

Some Perspective



1,000,000,000,000,000,000
= number of stopped Mu2e muons
= number of grains of sand on earth

Why so many?

- The new theories predict that the rate of $\mu N \rightarrow e N$ will occur at most

$$\frac{1}{1,000,000,000,000,000}$$

- This is a very very very very very very very small rate... so you need a lot of μ s to observe and measure it.

What is the status of Mu2e?

- Mu2e is currently in the Conceptual Design phase
- Mu2e hopes to break ground on the building within the next few years
- Mu2e plans to begin taking data by the end of this decade

The end is nigh

Summary

- High Energy Physics aims to answer two simple questions
 - What are the fundamental particles of nature?
 - What forces govern their interactions?
- We have a pretty good answer, the Standard Model

Summary

- High Energy Physics is at a crossroads
 - We know that the Standard Model is incomplete
 - We have lots of ideas about what a more complete model might look like
 - ... but we have no idea which is the right one

Summary

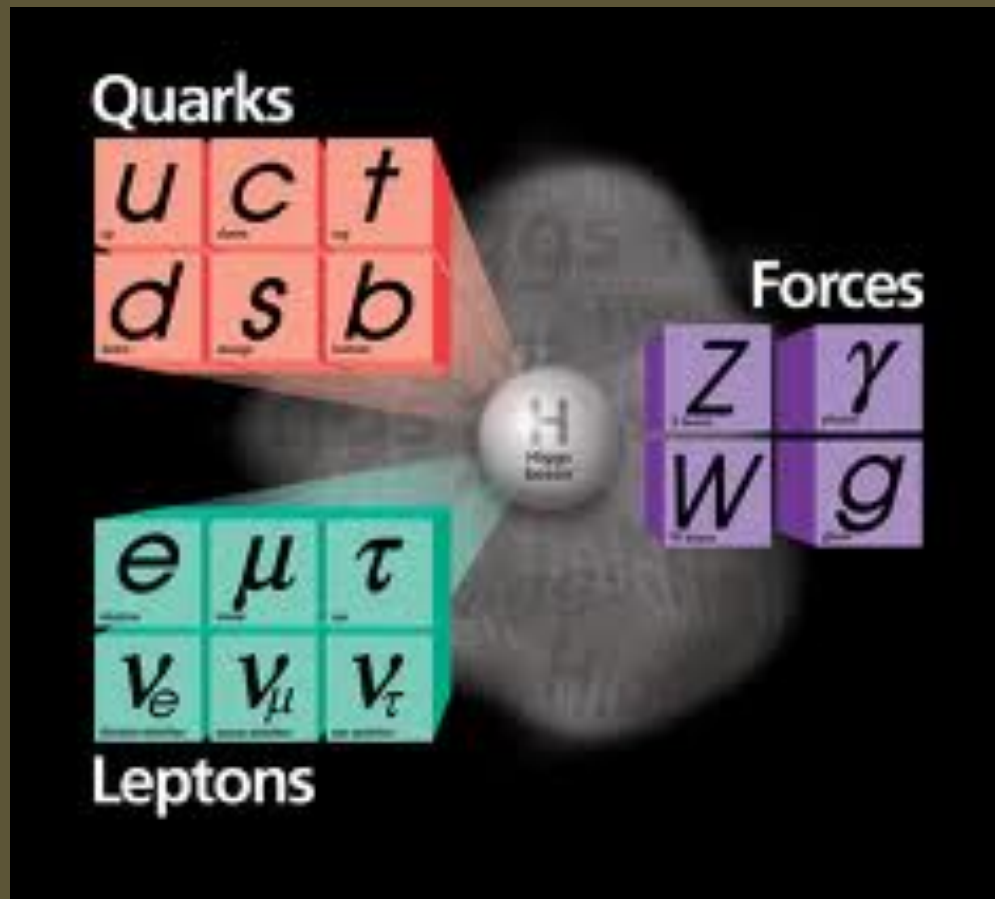


Fermilab's Mu2e experiment is important because it is designed to discover which direction is the right one

Thank You!

Additional Slides

Standard Model



Higgs boson gives mass to the quarks and leptons
(has not been verified experimentally)

Discoveries

What	When	Who	How
electron	1897	Thompson	cathodes
anti-electron	1932	Anderson	Cosmic rays
muon	1936	Anderson/Neddermeyer	Cosmic rays
tau	1975	Perl, et al	SLAC $ee \rightarrow em$
ν_e	1956	Cowan, Reines	$\nu + p \rightarrow e + n$
ν_μ	1962	Lederman, Steinberger, Schwartz	BNL, pion decay
ν_τ	2000	DONUT Collaboration	Fermilab
u, d, s quarks	1968	Breidenbach, Freidman, Kendall	Fermilab
c quark	1974	Richter, Ting	BNL, SLAC
b quark	1977	Lederman	Fermilab
t quark	1995	CDF, D0 Collaborations	Fermilab
photon	1905	Einstein	Theory work
gluon	1978	PLUTO Collaboration	DESY
W & Z bosons	1983	Rubbia, Van de Meer, et al	CERN